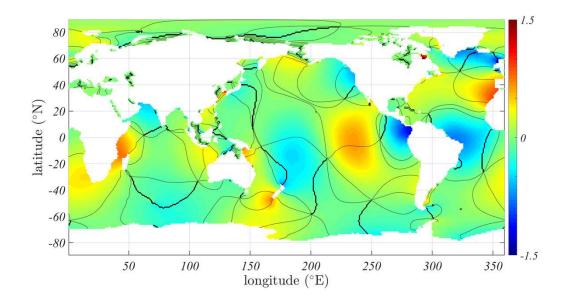
# <u>Lecture 6:</u> Shallow water dynamics and gravity waves

Callum J. Shakespeare

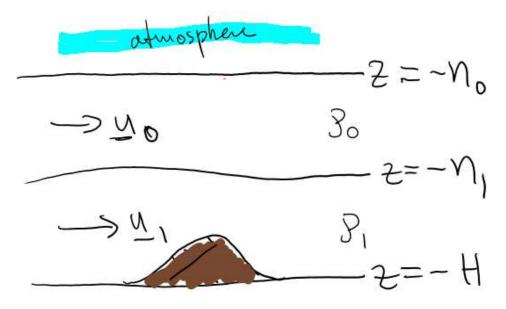
Fellow, Climate and Fluid Physics, ANU



#### Derivation of 2-layer shallow water equations

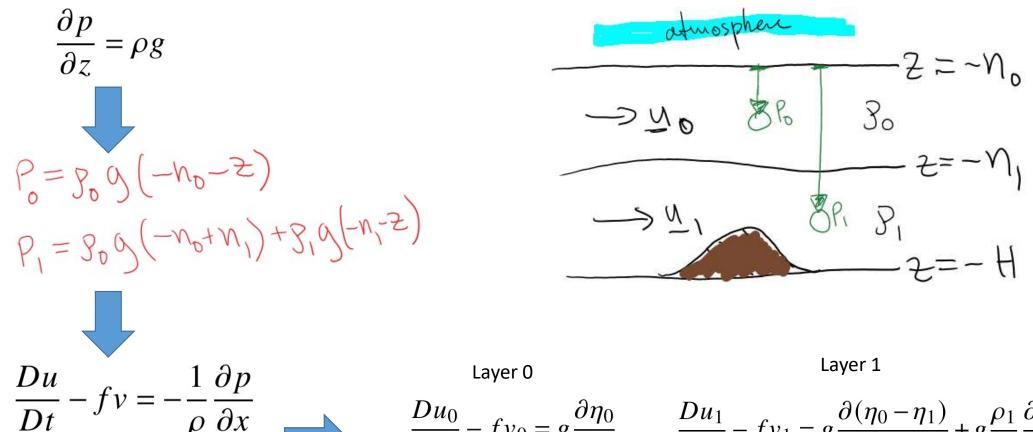
Hydrostatic momentum equations

$$\frac{Du}{Dt} - fv = -\frac{1}{\rho} \frac{\partial p}{\partial x}$$
$$\frac{Dv}{Dt} + fu = -\frac{1}{\rho} \frac{\partial p}{\partial y}$$
$$\frac{\partial p}{\partial z} = \rho g$$

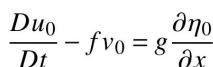


 $P_{1} = g_{0}g(-n_{0}-z)$   $P_{1} = g_{0}g(-n_{0}+n_{1})+g_{1}g(-n_{1}-z)$ 

#### Derivation of 2-layer shallow water equations



 $\frac{Dv}{Dt} + fu = -\frac{1}{2}\frac{\partial p}{\partial y}$ 



 $\frac{Du_0}{Dt} - fv_0 = g\frac{\partial\eta_0}{\partial x} \qquad \frac{Du_1}{Dt} - fv_1 = g\frac{\partial(\eta_0 - \eta_1)}{\partial x} + g\frac{\rho_1}{\rho_0}\frac{\partial\eta_1}{\partial x}$  $=g\frac{\partial\eta_0}{\partial x} + g\frac{\rho_1 - \rho_0}{\rho_0}\frac{\partial\eta_1}{\partial x}$ 

#### Derivation of 2-layer shallow water equations

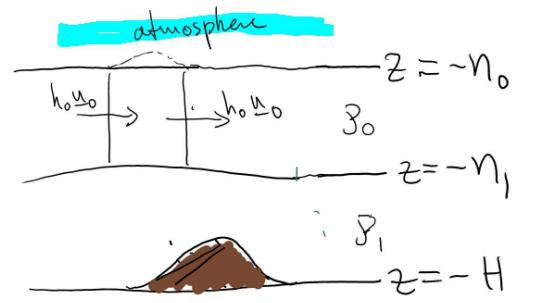
Conservation of mass

$$h_{o} = N_{1} - N_{o}$$

$$\frac{\partial h_{o}}{\partial t} = -\nabla \cdot (\text{volume flux})$$

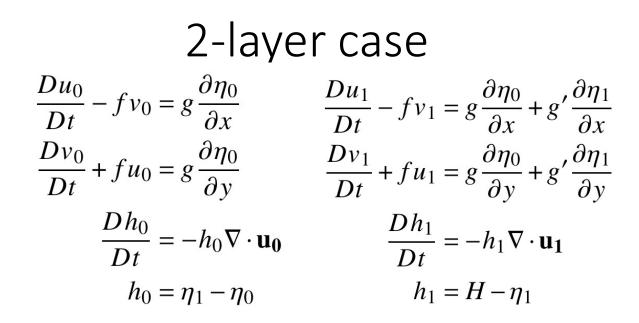
$$= -\nabla \cdot (h_{o} \underline{u}_{o})$$

$$= -h_{o} \nabla \cdot \underline{u}_{o} - \underline{u}_{o} \nabla h$$



$$\frac{Dh_0}{Dt} = -h_0 \nabla \cdot \mathbf{u_0}$$

Layer 1 similar but  $h_1 = H - \eta_1$ 



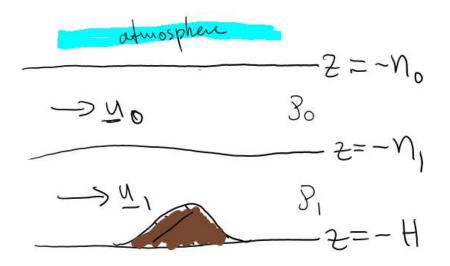
1-layer case  

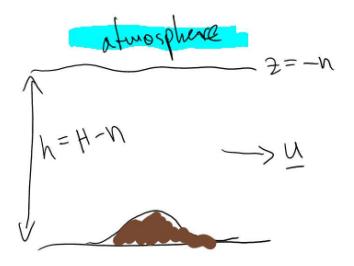
$$\frac{Du}{Dt} - fv = g \frac{\partial \eta}{\partial x}$$

$$\frac{Dv}{Dt} + fu = g \frac{\partial \eta}{\partial y}$$

$$\frac{Dh}{Dt} = -h\nabla \cdot \mathbf{u}$$

$$h = H - \eta$$



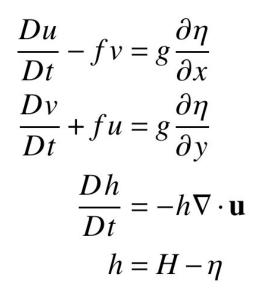


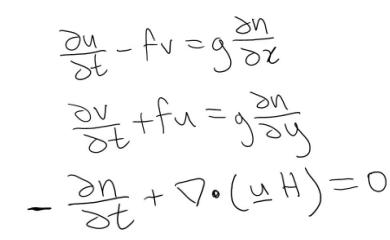
#### Linearisation and wave solutions

Non-linear equations

Linearised equations

Assumptions



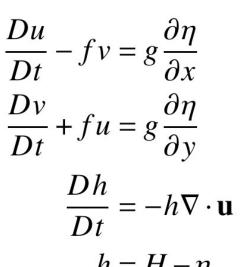


 $\left(\underline{N}\cdot\Delta<<\frac{2}{3}=2,\frac{M}{M}<<1\right)$ 

 $(\eta < \leq \#)$ 

#### Linearisation

Non-linear equations



Linearised equations

 $\frac{\partial u}{\partial t} - fv = g \frac{\partial n}{\partial z}$ 

 $\frac{\partial v}{\partial t} + fu = g \frac{\partial n}{\partial u}$ 

 $-\frac{\partial n}{\partial t}+\nabla \cdot (\Psi H)=0$ 

Assumptions

 $\left(\underline{N}\cdot\Delta << \frac{2}{9} = \frac{1}{N} << 1\right)$ 

 $(\eta < = \#)$ 

$$\begin{array}{c} u = u \\ let \\ \downarrow = u \\ \hline \\ \hline \\ \hline \\ \hline \\ \hline \\ \hline \\ \hline \\ \\ \hline \\ \\ \hline \\ \\ \hline \\ \hline \\ \\ \hline \\ \\ \hline \\ \hline \\ \\ \\ \hline \\ \hline \\ \hline \\ \\ \hline \\ \hline \\ \\ \hline \\ \hline \\ \hline \\ \hline \\ \hline \\ \\ \hline \hline \\ \hline \\ \hline \\ \hline \hline$$

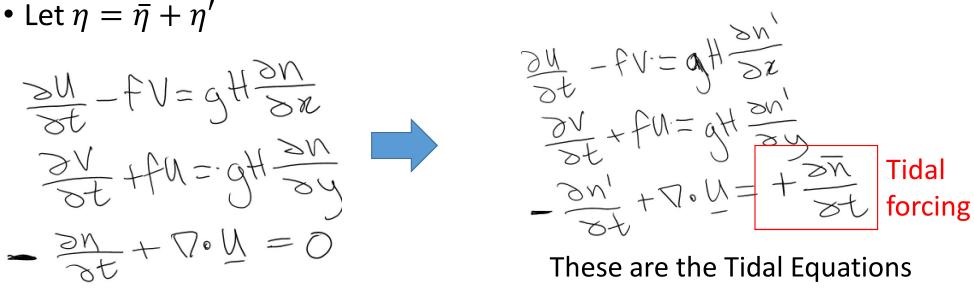
# Wave solutions $\frac{\partial U}{\partial t} - fV = gH \frac{\partial n}{\partial x} \qquad \text{let } d = \hat{\psi} e^{i(kx+ly-wt)}$ $\frac{\partial V}{\partial t} + fu = igH \frac{\partial n}{\partial y} \qquad \text{let } d = \hat{\psi} e^{i(kx+ly-wt)}$ $-iw\hat{\psi} - f\hat{V} = ikgH\hat{\eta}$ $-iw\hat{\psi} + f\hat{\eta} = ilgH\hat{\eta}$ $+iw\hat{\eta} + ik\hat{\eta} + il\hat{\psi} = 0$ **Dispersion Relation**

$$\omega^{2} = f^{2} + (k^{2} + l^{2})gH$$
$$\omega = \pm \sqrt{f^{2} + K^{2}gH}$$

Barotropic gravity waves (BGWs)

#### How do we generate BGWs?

- Full depth (barotropic) forcing = body force
- E.g. horizontal variations in gravity due to the moon = tidal forces!
- Usually represented in terms of an equilibrium tide = height perturbation that would exist if no flow:  $\bar{\eta} = A \cos(\omega_t t + 2\phi)$
- Flow is associated with departures from  $ar\eta$



or if on a sphere "Laplace Tidal Equations"

• Forced at a particular frequency = 12.4 hour period

$$\omega = \pm \sqrt{f^2 + K^2 g H}$$

• 
$$f \sim \frac{10^{-4}}{s}$$
,  $\omega \sim 1.4 \times \frac{10^{-4}}{s}$ ,  $g \sim 10 \frac{m}{s^2}$ ,  $H \sim 4000 m$ 

- What is the value of K?
- $\frac{1}{K} \sim 2000 \text{ km}$

• Forced at a particular frequency = 12.4 hour period

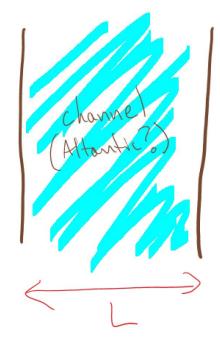
$$\omega = \pm \sqrt{f^2 + K^2 g H}$$

• 
$$f \sim \frac{10^{-4}}{s}$$
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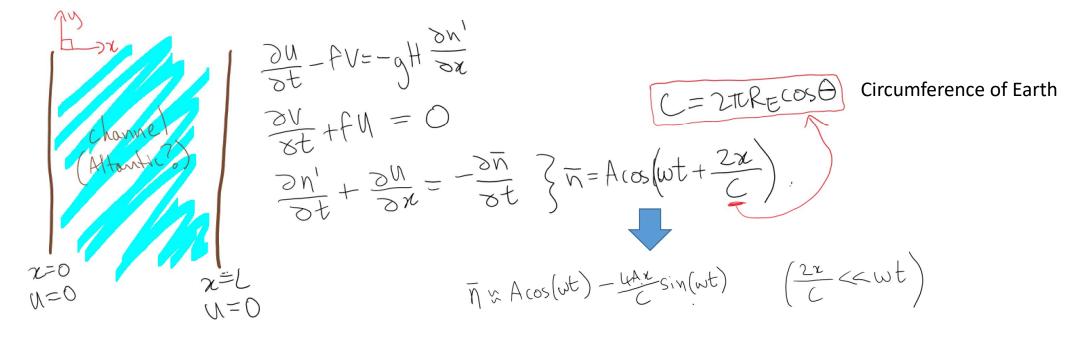
- What is the value of K?
- $\frac{1}{K} \sim 2000 \text{ km}$

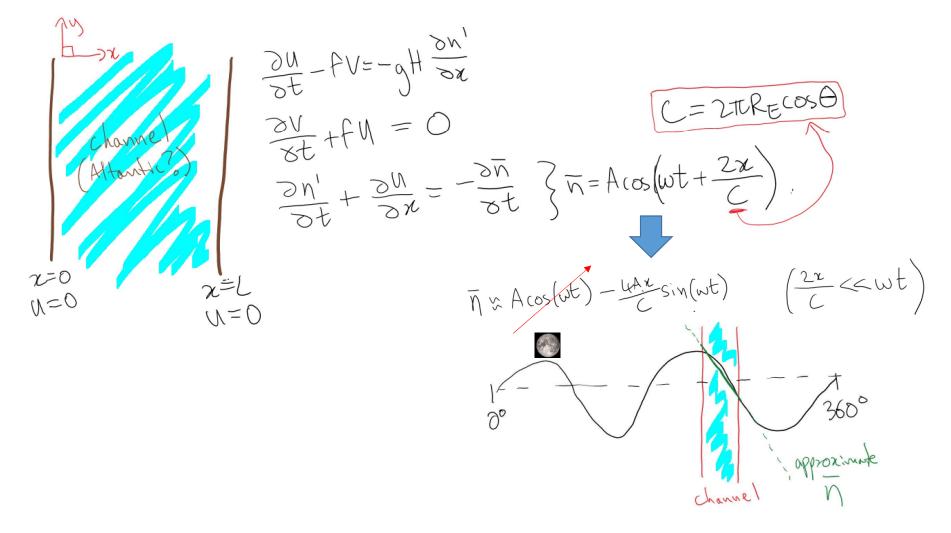




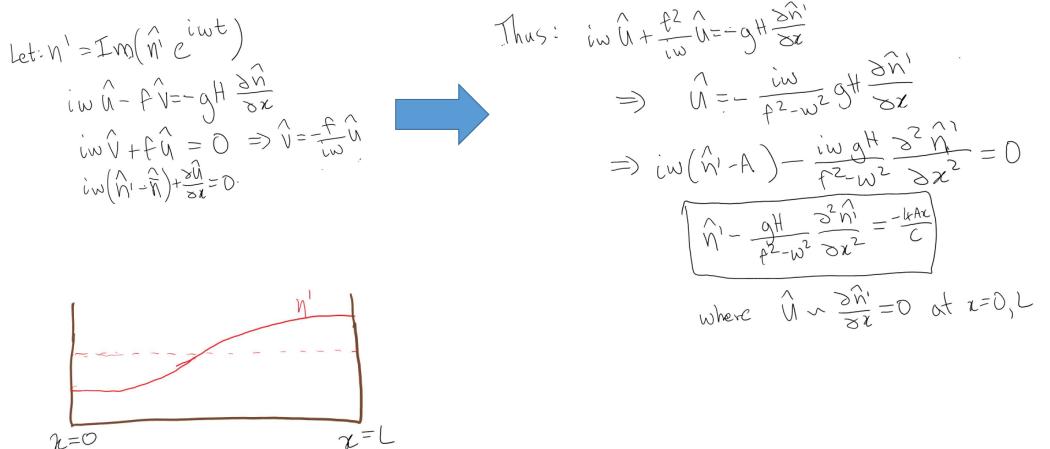


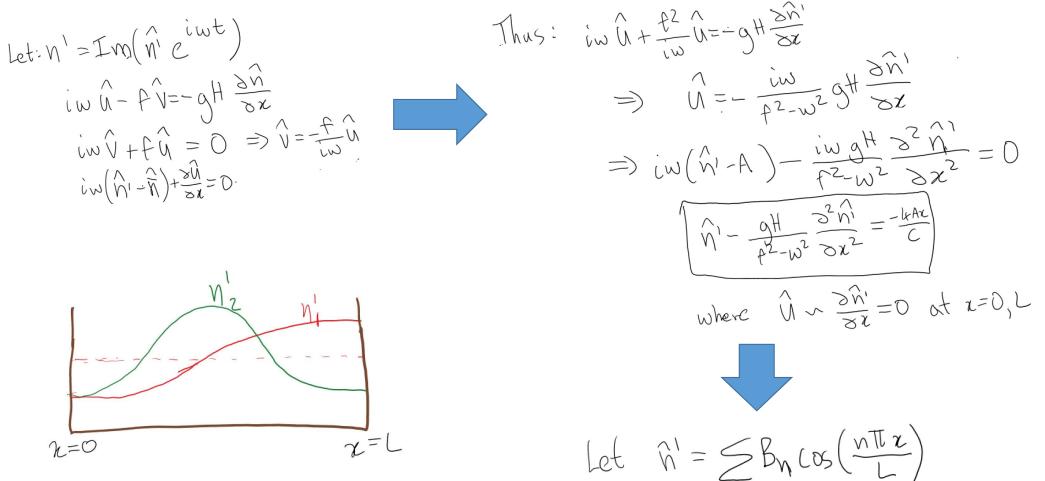
- Consider a simplified channel
- Infinitely long in y
- Width L in x





$$\begin{aligned} \text{Let: } \mathbf{N}' = \mathbf{I} \cdot \mathbf{M} \left( \hat{\mathbf{n}}' \in \overset{\text{i} \cdot \text{wt}}{\mathbf{n}} \right) \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) = -g H \frac{\partial \hat{\mathbf{n}}}{\partial \mathbf{x}} \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) = -g H \frac{\partial \hat{\mathbf{n}}}{\partial \mathbf{x}} \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) = -g H \frac{\partial \hat{\mathbf{n}}}{\partial \mathbf{x}} \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) - \frac{i \cdot \omega}{A^2 - \omega^2} \frac{\partial H}{\partial \mathbf{x}^2} = 0 \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) - \frac{i \cdot \omega}{A^2 - \omega^2} \frac{\partial H}{\partial \mathbf{x}^2} = 0 \\ & \text{i} \cdot \omega \left( \hat{\mathbf{n}} - A \right) - \frac{\partial H}{A^2 - \omega^2} \frac{\partial^2 \hat{\mathbf{n}}'}{\partial \mathbf{x}^2} = 0 \\ & \text{where } \hat{\mathbf{n}} \cdot \frac{\partial \hat{\mathbf{n}}'}{\partial \mathbf{x}^2} = 0 \text{ at } \mathbf{z} = 0, \end{aligned}$$





Let 
$$\hat{n}' = \leq B_{n} \cos\left(\frac{n \operatorname{Tr} z}{L}\right)$$
  $\left(\hat{n}' - \frac{\alpha H}{P^{2} - \omega^{2}} \frac{\partial^{2} \hat{n}'}{\partial x^{2}} = -\frac{4Az}{C}\right)$ 

$$= \frac{gH}{n} \cos\left(\frac{n\pi z}{L}\right) - \frac{gH}{f^2 - w^2} \frac{\partial^2}{\partial x^2} B_n \cos\left(\frac{n\pi z}{L}\right) = -\frac{4A}{C} \sum_n C_n \cos\left(\frac{n\pi z}{L}\right)$$
where  $\chi = \sum_n C_n \cos\left(\frac{n\pi z}{L}\right)$ 

Let 
$$\hat{n}' = \leq B_n \cos\left(\frac{n\pi r}{L}\right)$$
  
 $\left[\hat{n}' - \frac{\alpha H}{p^2 - \omega^2} \frac{\partial^2 n}{\partial x^2} = -\frac{4\alpha r}{C}\right]$   
 $\leq B_n \cos\left(\frac{n\pi r}{L}\right) - \frac{\alpha H}{f^2 - \omega^2} \frac{\partial^2}{\partial x^2} B_n \cos\left(\frac{n\pi r}{L}\right) = -\frac{4\alpha r}{C} \leq C_n \cos\left(\frac{n\pi r}{L}\right) \times \cos\left(\frac{m\pi r}{L}\right) \int_0^L dx$   
where  $x = \leq C_n \cos\left(\frac{n\pi r}{L}\right) \times \cos\left(\frac{m\pi r}{L}\right) \int_0^L dx$ 

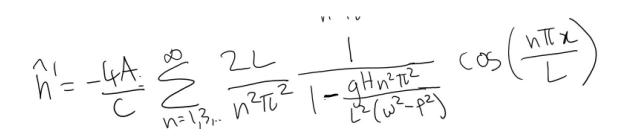
$$et \quad \hat{n}' = \leq B_{n} \cos\left(\frac{n \operatorname{Tr} z}{L}\right) \qquad \left(\hat{n}' - \frac{q H}{p^{2} - w^{2}} \frac{3^{2} n^{2}}{3 x^{2}} = -\frac{\mu A x}{C}\right)$$

$$= \frac{4 A}{C} \leq \frac{1}{n} \cos\left(\frac{n \operatorname{Tr} z}{L}\right) - \frac{q H}{p^{2} - w^{2}} \frac{3^{2}}{3 x^{2}} B_{n} \cos\left(\frac{n \operatorname{Tr} z}{L}\right) = -\frac{4 A}{C} \leq \frac{1}{n} \cos\left(\frac{n \operatorname{Tr} x}{L}\right) \times \cos\left(\frac{m \operatorname{Tr} x}{L}\right) \int_{0}^{L} dz$$

$$= B_{n} \cos\left(\frac{n \operatorname{Tr} z}{L}\right) - \frac{q H}{p^{2} - w^{2}} \frac{3^{2}}{3 x^{2}} B_{n} \cos\left(\frac{n \operatorname{Tr} x}{L}\right) = -\frac{4 A}{C} \leq \frac{1}{n} \cos\left(\frac{n \operatorname{Tr} x}{L}\right) \times \cos\left(\frac{m \operatorname{Tr} x}{L}\right) \int_{0}^{L} dz$$

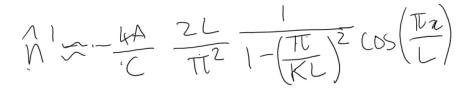
$$= \sum_{n} \cos\left(\frac{n \operatorname{Tr} x}{L}\right) \times \cos\left(\frac{m \operatorname{Tr} x}{L}\right) \int_{0}^{L} dz$$

$$C_{n} = \int_{0}^{L} \cos\left(\frac{n\pi n}{L}\right) \chi d\chi$$
  
$$= \int_{0}^{L} \cos^{2}\left(\frac{n\pi n}{L}\right) d\chi$$
  
$$= \int_{0}^{L} \cos^{2}\left(\frac{n\pi n}{L}\right) d\chi$$
  
$$= \int_{0}^{0} n \text{ even}$$
  
$$\frac{2L}{n^{2}\pi^{2}} n \text{ odd}$$



Mode n=3 is ~10% of mode n=1... so ignore it.

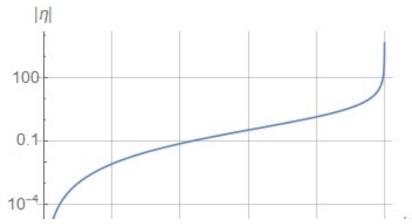
How is the size of my ocean basin related to the amplitude of the tide????

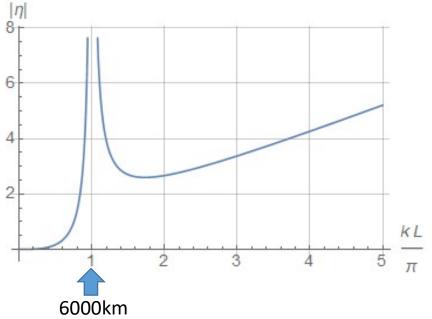


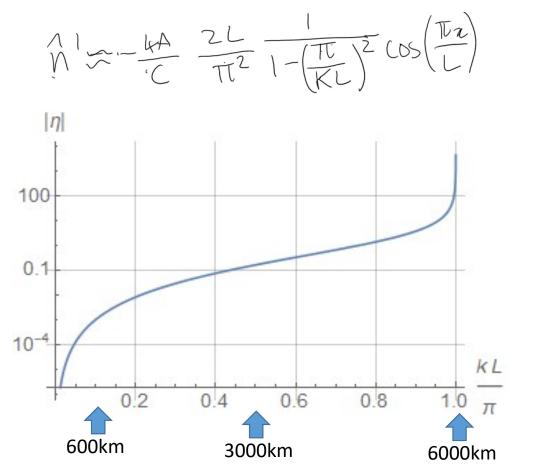
Here we have used the previous definition

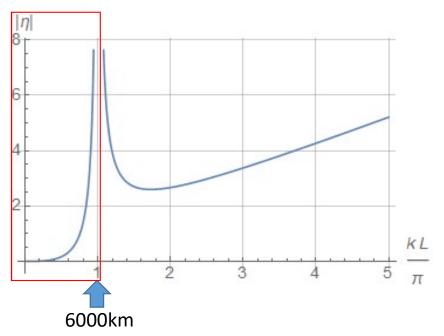
$$\omega = \pm \sqrt{f^2 + K^2 g H}$$

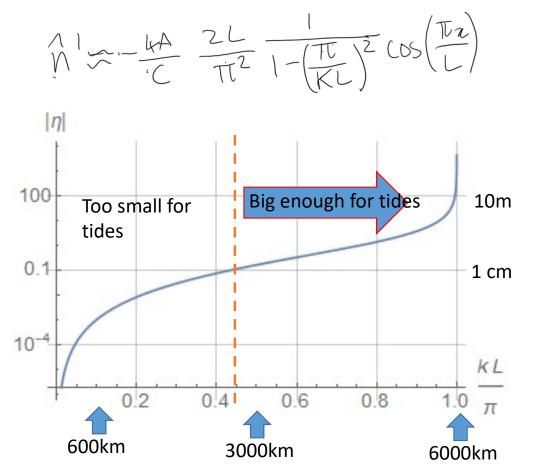
where  $\frac{1}{K} \sim 2000$  km

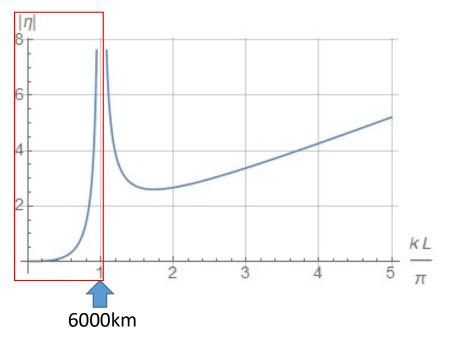






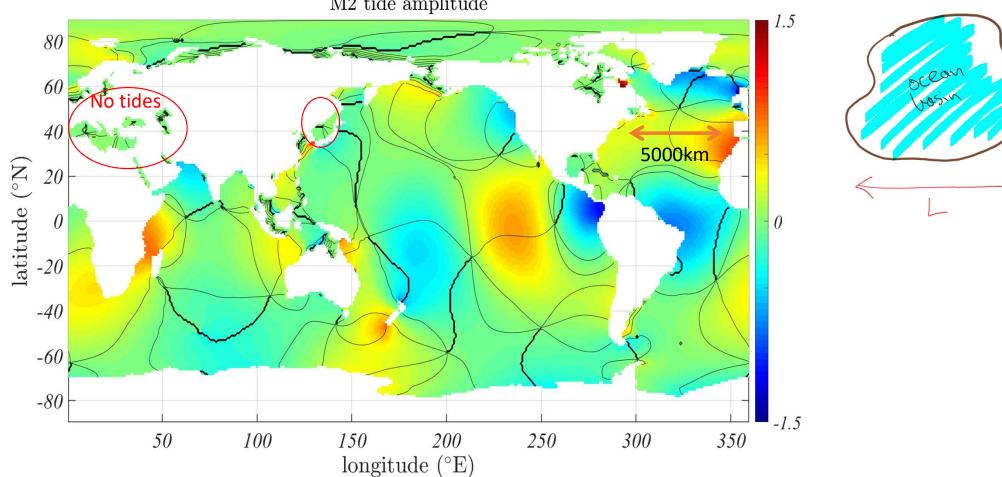






#### Comparison

#### How is the size of my ocean basin related to the amplitude of the tide????



M2 tide amplitude

#### Comparison

#### How is the size of my ocean basin related to the amplitude of the tide????

0.5 1 1.5 0 **TPXO9:M2** 400 00 Hudson Bay 270 m deep max L>500km for tides 40°5 @Egbert & Erofeeva, 2010 Grant NNX17AJ34G 120°E College of Earth, Ocean, 60°E 00 120°W 60°W 180°W 

https://www.tpxo.net/global

#### Comparison with Atlantic basin over time

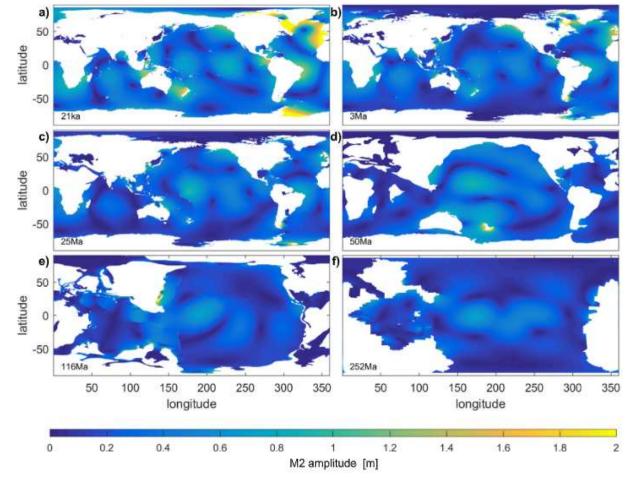


Fig. 2. Shown are the M2 tidal amplitudes for the LGM (a), Pliocene (b), Miocene (c), Eocene (d), Cretaceous (e) and Permian-Triassic (f).

From Green et al., 2017

#### Comparison with Atlantic basin over time

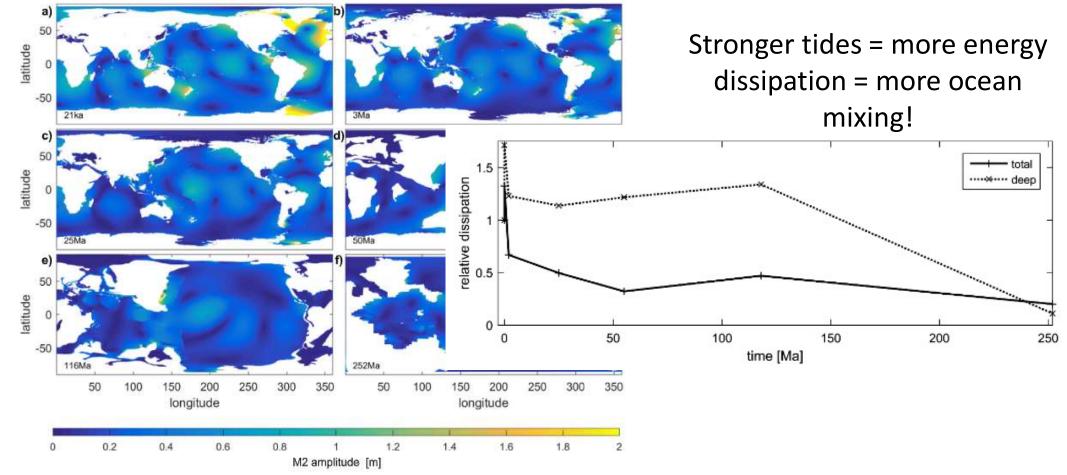
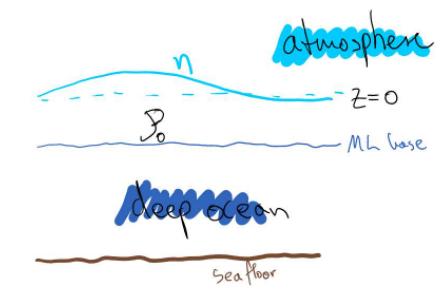


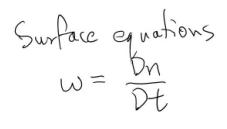
Fig. 2. Shown are the M2 tidal amplitudes for the LGM (a), Pliocene (b), Miocene (c), Eocene (d), Cretaceous (e) and Permian-Triassic (f).

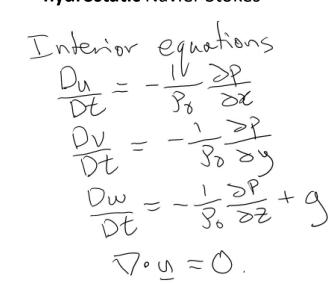
From Green et al., 2017

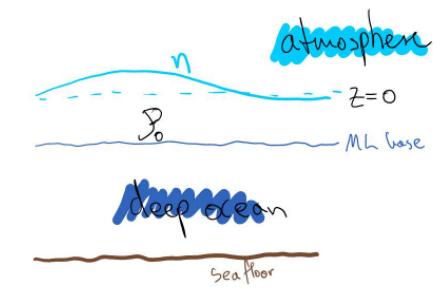


Kinematic boundary condition

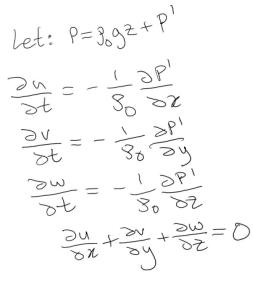
Uniform density, **nonhydrostatic** Navier Stokes

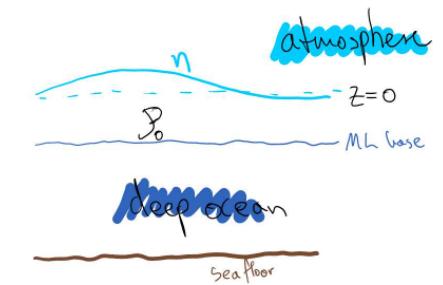


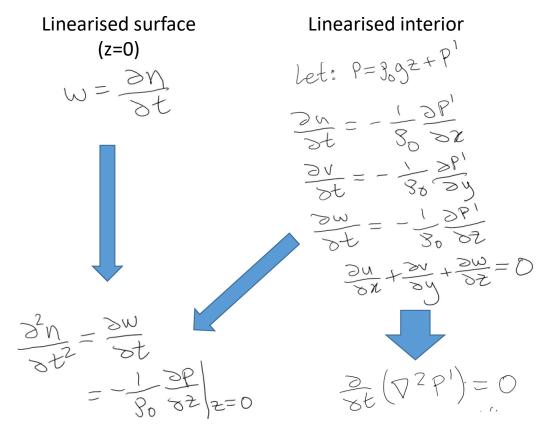


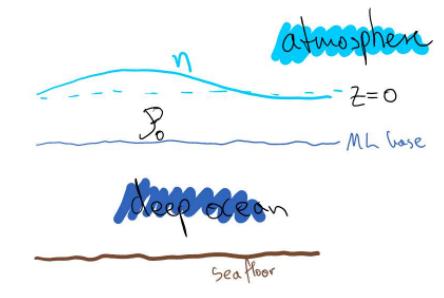


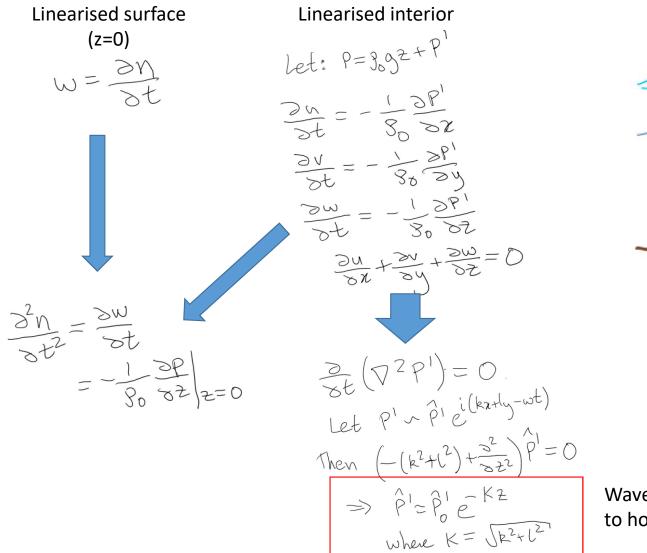
Linearised surface (z=0)  $w = \frac{\partial N}{\partial t}$  Linearised interior

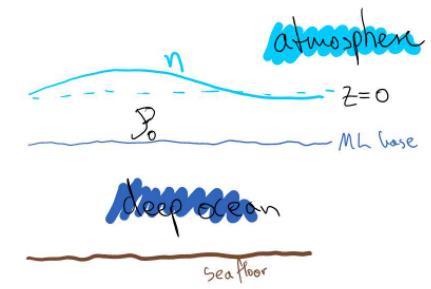




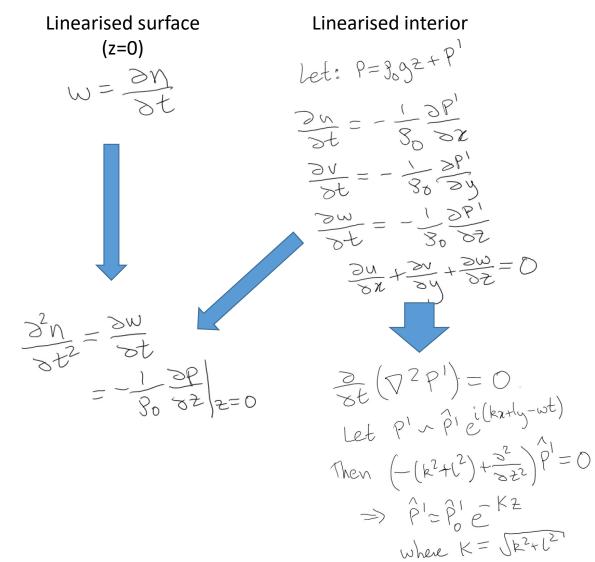


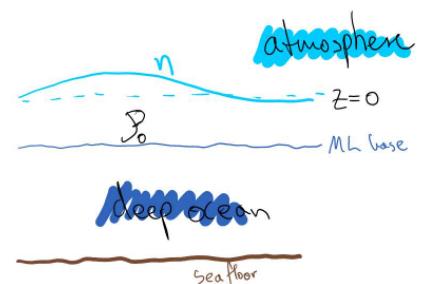




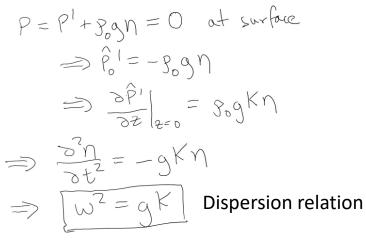


Wave amplitude decays with depth, according to horizontal wavelength





At the surface the total pressure must be constant to match atmospheric pressure, thus



## Surface gravity waves (SGWs)

$$\omega = \sqrt{gK}$$
$$P \sim \rho_0 g \eta e^{-Kz}$$



#### SGWs

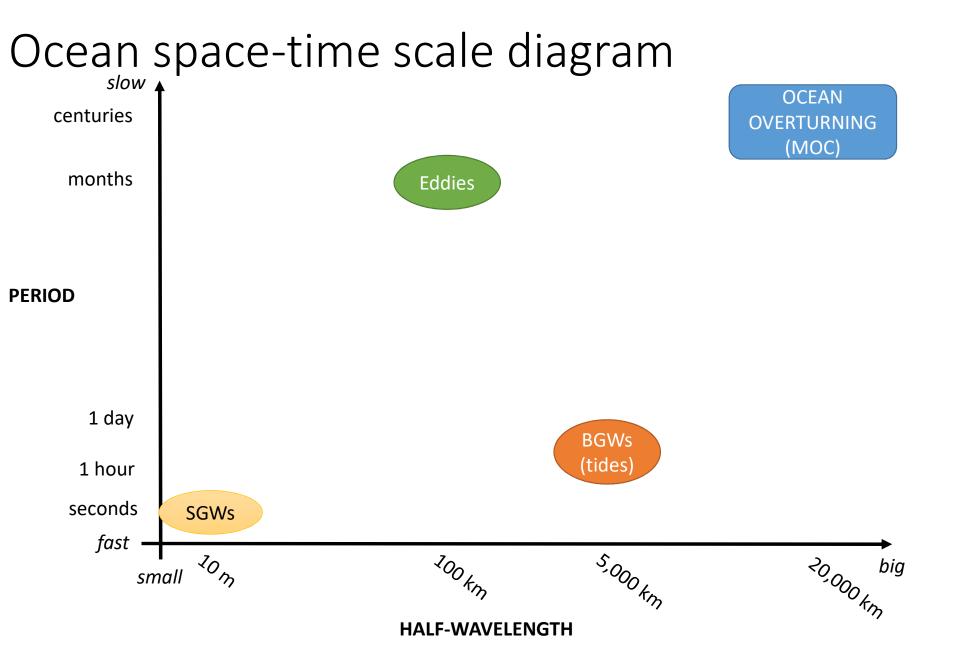
$$\omega = \pm \sqrt{gK}$$

- Generated by perturbations at the surface (wind!)
- $\frac{2\pi}{\omega} \sim 3 \ seconds \Rightarrow \frac{2\pi}{k} \sim 14 \ m$
- Small spatial scales
- Non-hydrostatic = large vertical acceleration
- Shallow, surface intensified motion

#### BGWs

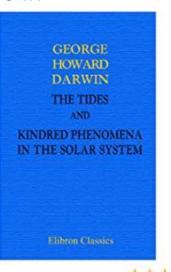
$$\omega = \pm \sqrt{f^2 + K^2 g H}$$

- Generated by full depth/body force perturbations (tidal forces)
- $\frac{2\pi}{\omega} \sim 12.4 \ hrs \Rightarrow \frac{2\pi}{k} \sim 10,000 \ km$
- Large spatial scale
- Hydrostatic = slow vertical acceleration
- Full-depth barotropic motion



## Some references

- Tides and modelling thereof:
  - Arbic, Brian K., et al. "A primer on global internal tide and internal gravity wave continuum modeling in HYCOM and MITgcm." *New frontiers in operational oceanography* (2018).
- Tides, seiches; history of observation ------
- Surface gravity waves theory
  - Acheson, David J. "Elementary fluid dynamics." (1991)



- Paleo-tides
  - Green. "Ocean tides and resonance." Ocean Dynamics 60.5 (2010): 1243-1253.
  - Green, J. A. M., et al. "Explicitly modelled deep-time tidal dissipation and its implication for Lunar history." *Earth and Planetary Science Letters* 461 (2017): 46-53.

Spherical coordinates: water world  

$$\nabla_{h} \cdot \mathbf{A} = \frac{1}{r \cos \theta} \frac{\partial}{\partial \phi} A_{\phi} + \frac{1}{r \cos \theta} \frac{\partial}{\partial \theta} A_{\theta} \cos \theta$$

$$\nabla_{h} A = \left(\frac{1}{r \cos \theta} \frac{\partial A}{\partial \phi}, \frac{1}{r} \frac{\partial A}{\partial \theta}\right)$$

$$f = 2\Omega \sin \theta$$

